



Nonidentical particle correlations

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Faculty of Physics

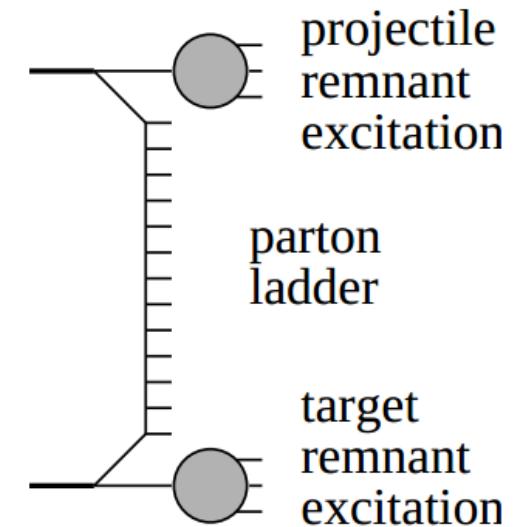
Outline

- EPOS model
 - Spectra p_T from EPOS and STAR²
 - Four-vector of space and momentum
- Correlation functions
 - Space-time asymmetry
 - STAR data
 - Results from EPOS
 - Centrality dependence
 - Fits
 - Sizes calculated from correlation functions
 - STAR and EPOS data for 39 GeV
- Summary
- Annotation

EPOS model¹

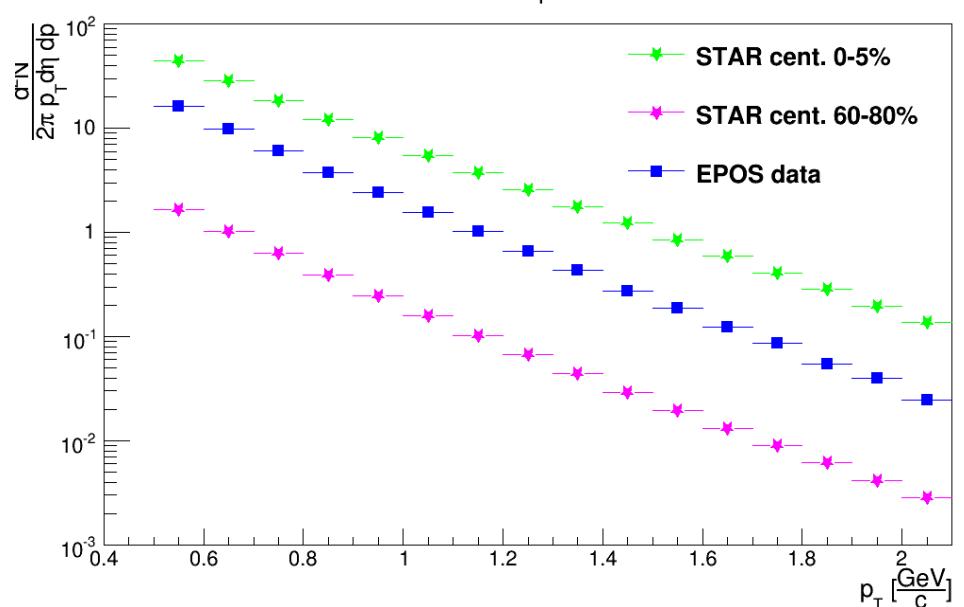
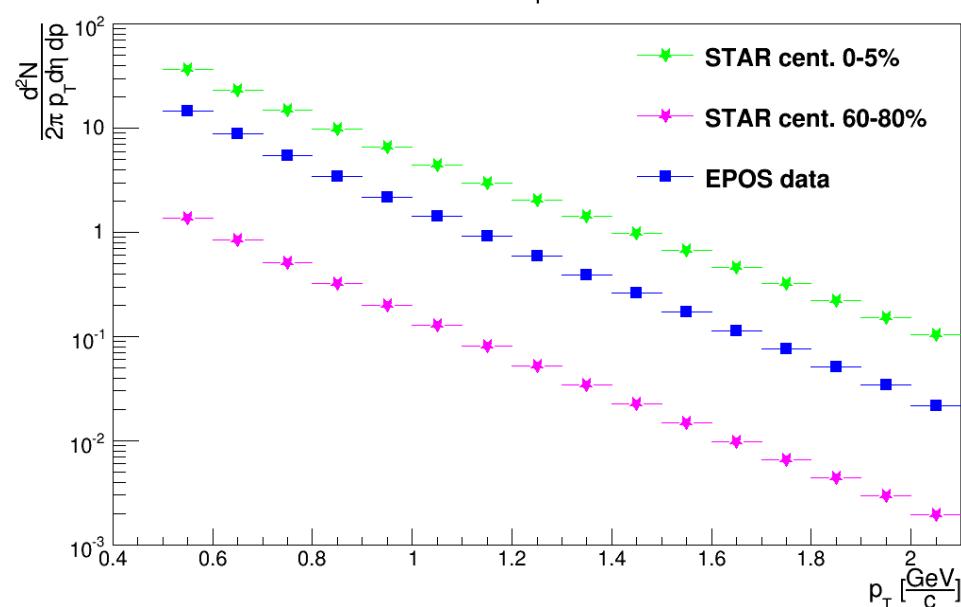
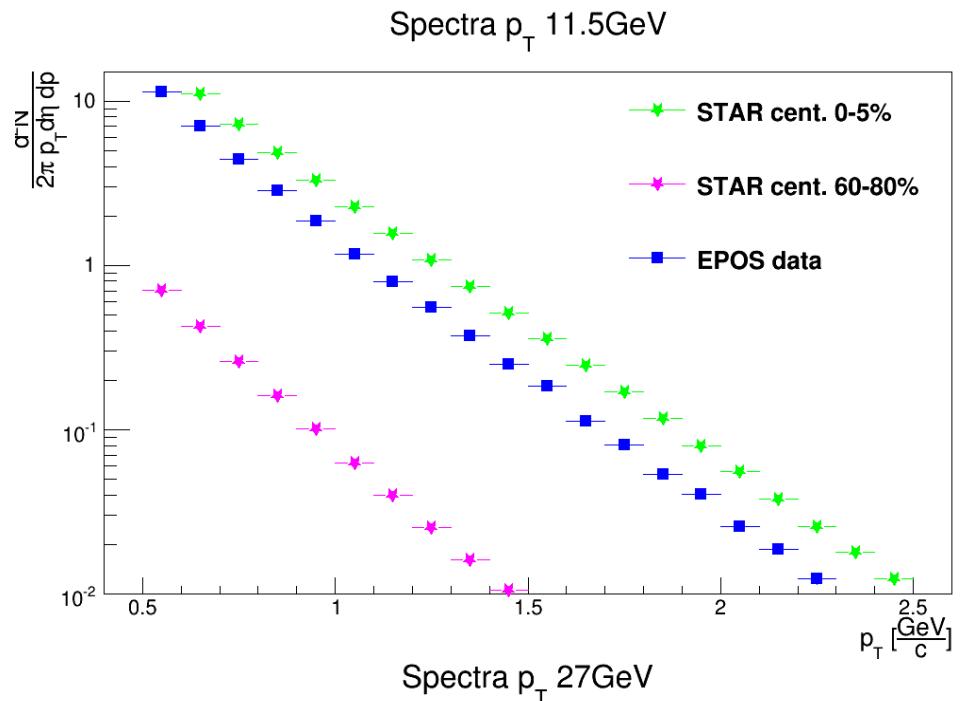
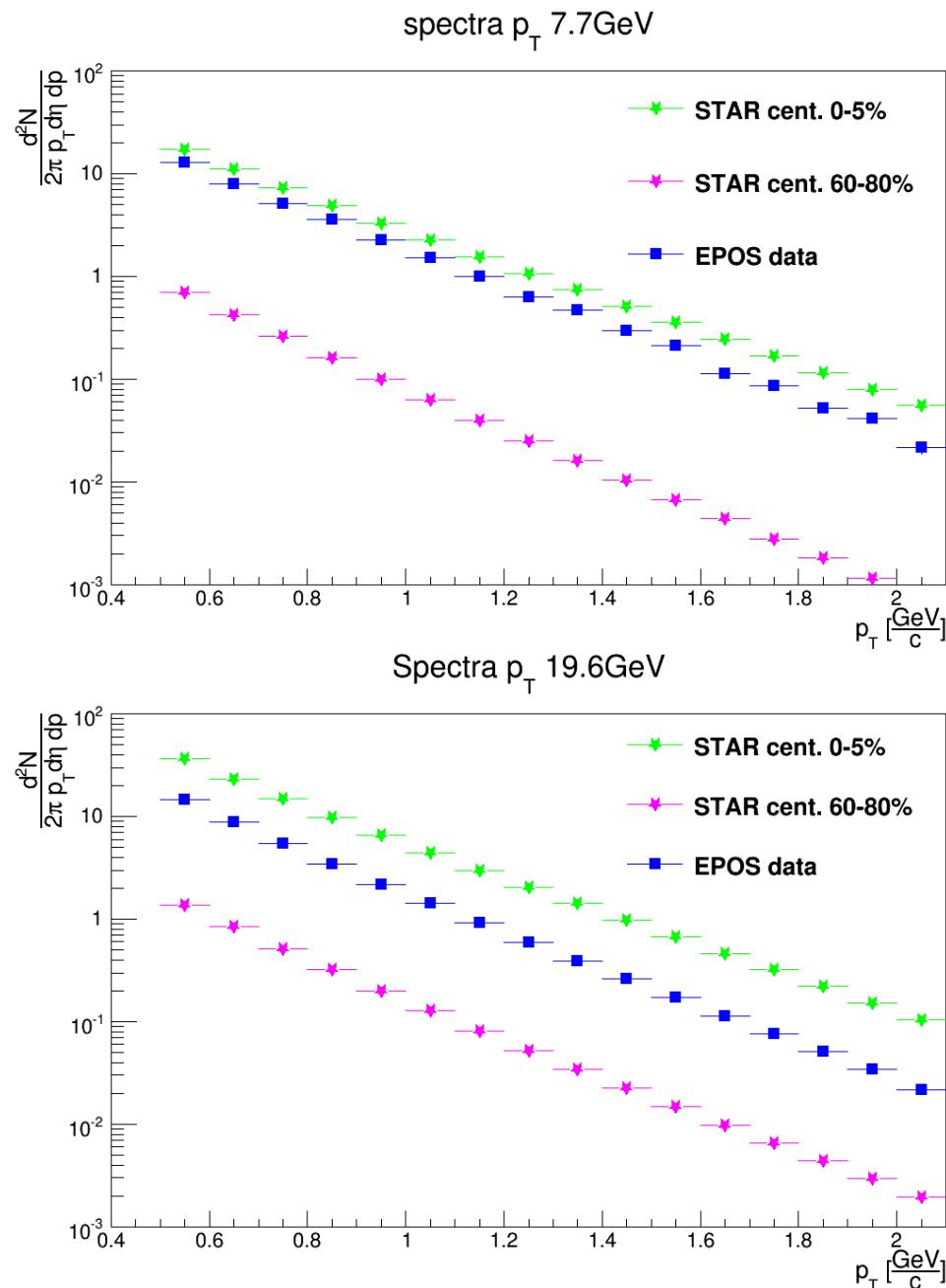
EPOS is a parton model, with many binary parton-parton interactions, where each one creating a parton ladder.

- Energy-sharing : for cross section calculation and particle production
- Parton Multiple scattering
- Outshell remnants
- Screening and shadowing via unitarization and splitting
- Collective effects for dense systems (LHC energies)
- We do not have the hydrodynamics for low RHIC energies.



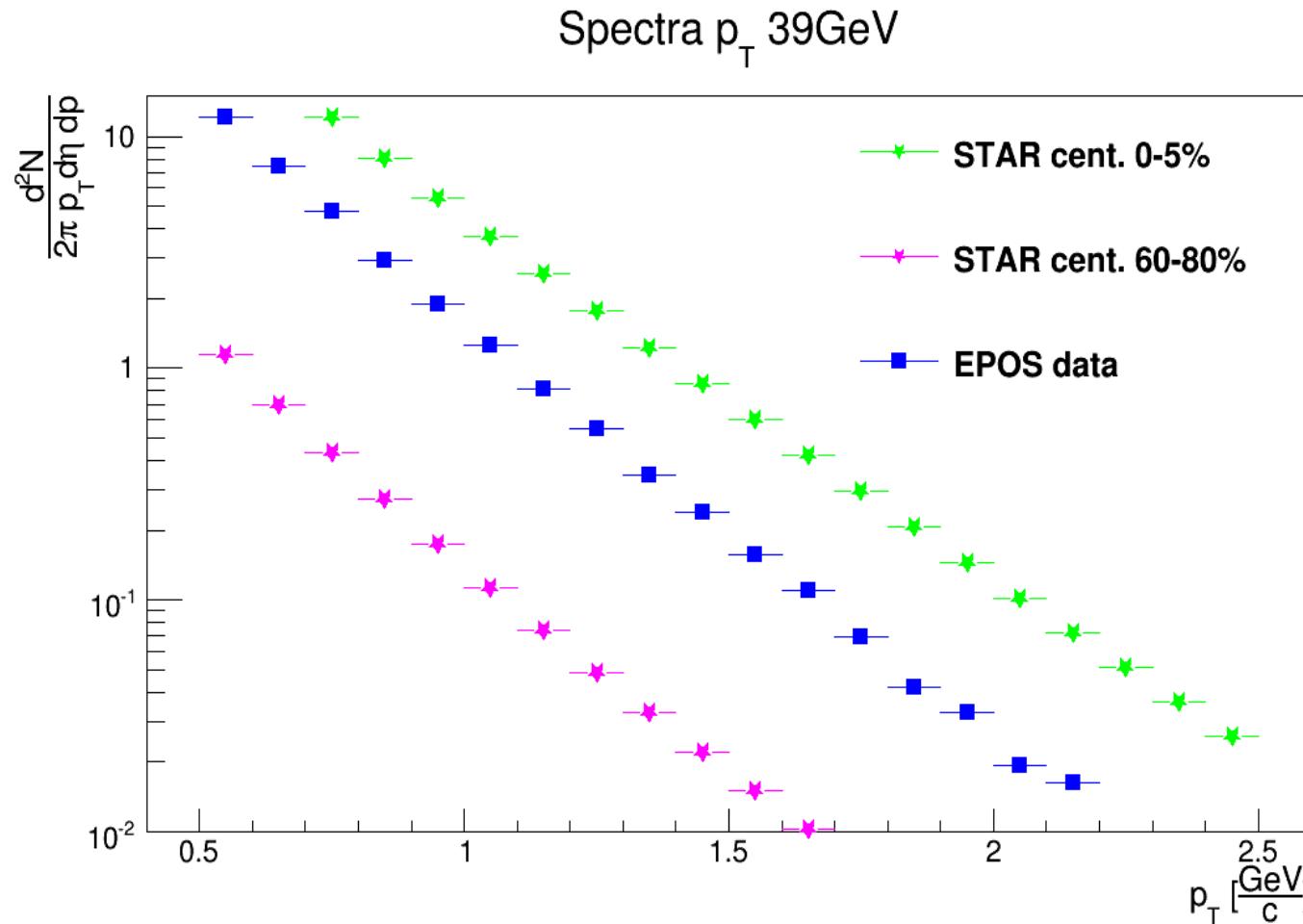
Author of the EPOS model is Klaus Werner from SUBATECH, University of Nantes – IN2P3/CNRS– EMN, Nantes, France.

Spectra p_T from EPOS and STAR²



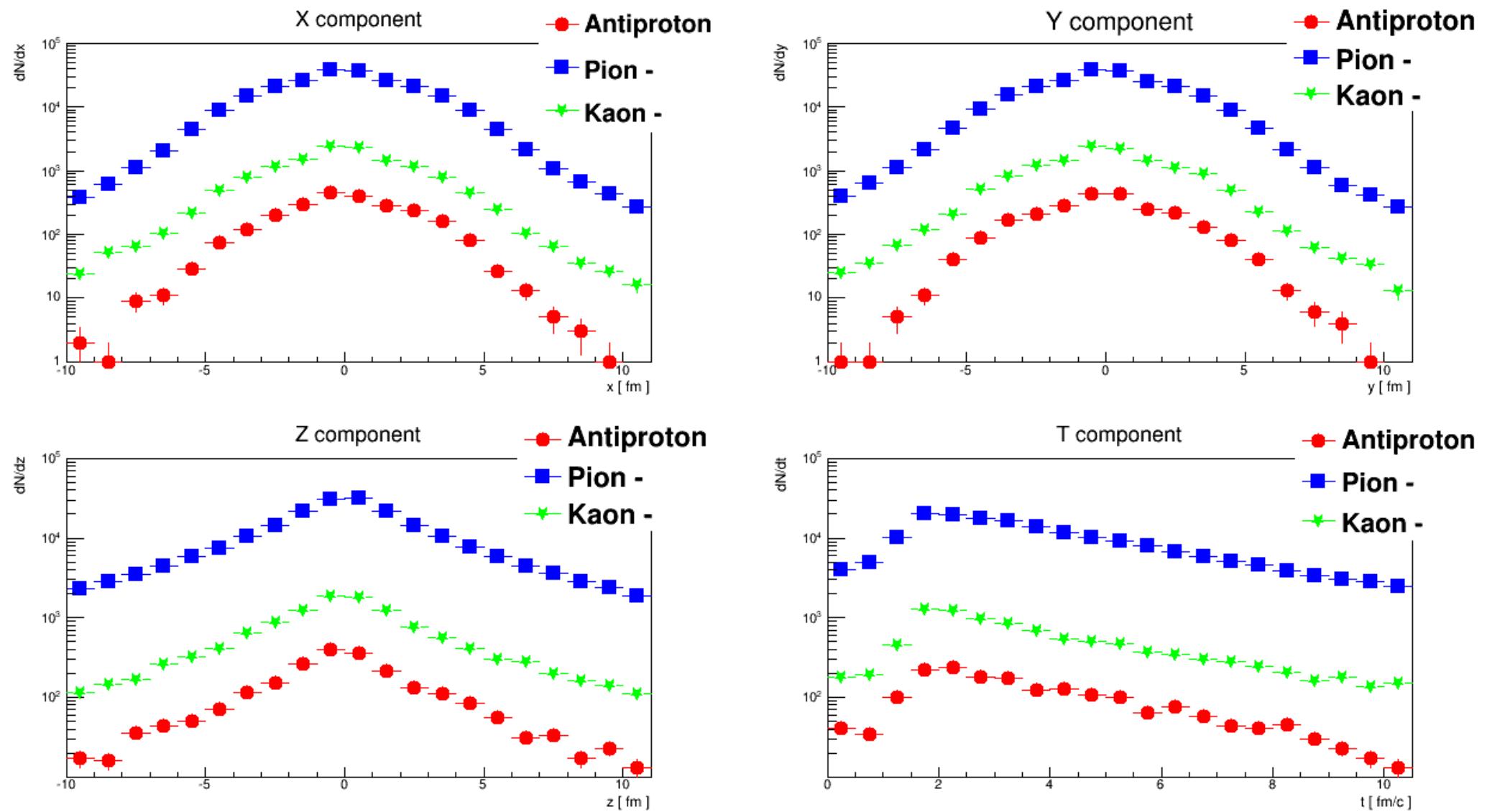
Spectra p_T from EPOS and STAR²

- Transverse momentum spectra results for charged particles from EPOS model are between STAR p_T spectra for the most central collisions (0-5%) and peripheral collisions (60-80%).



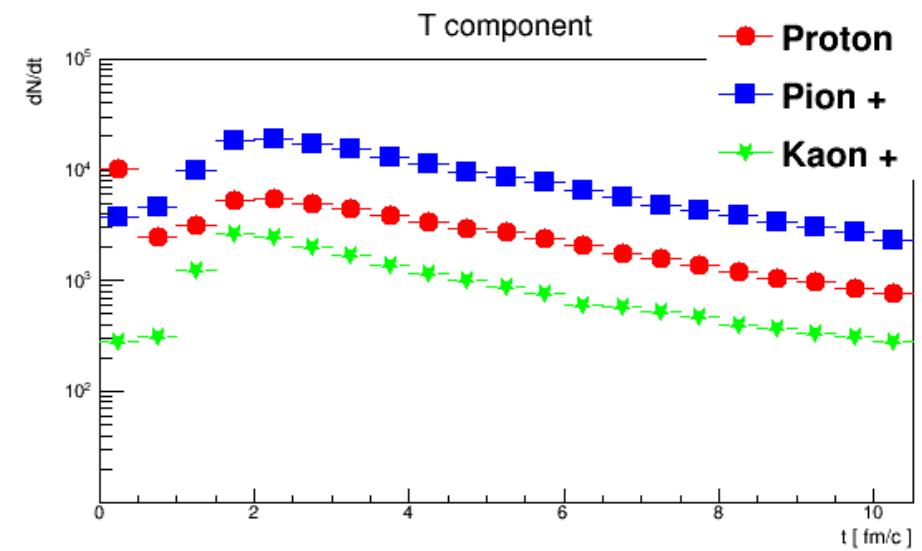
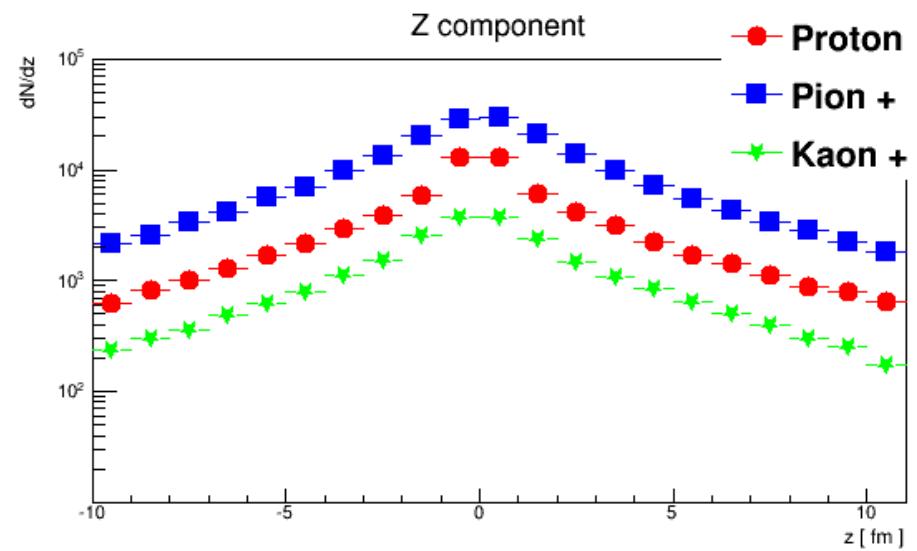
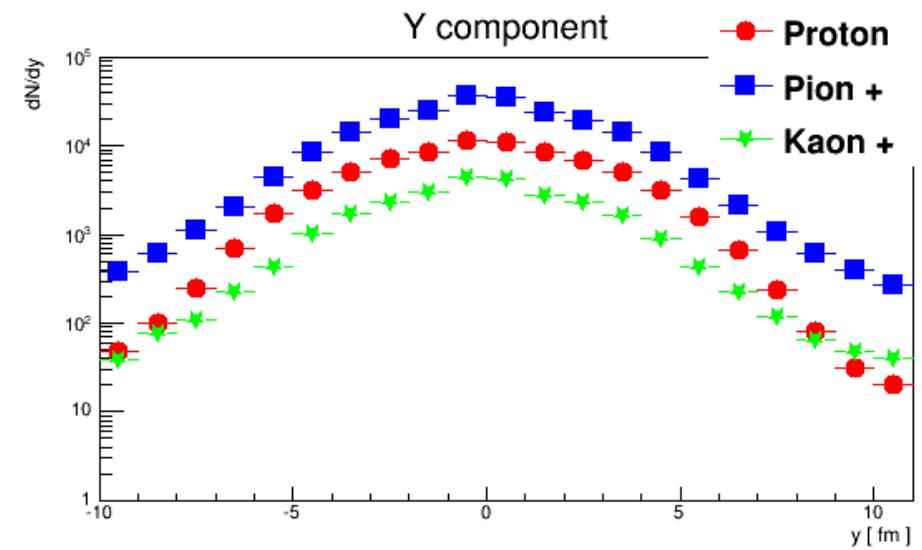
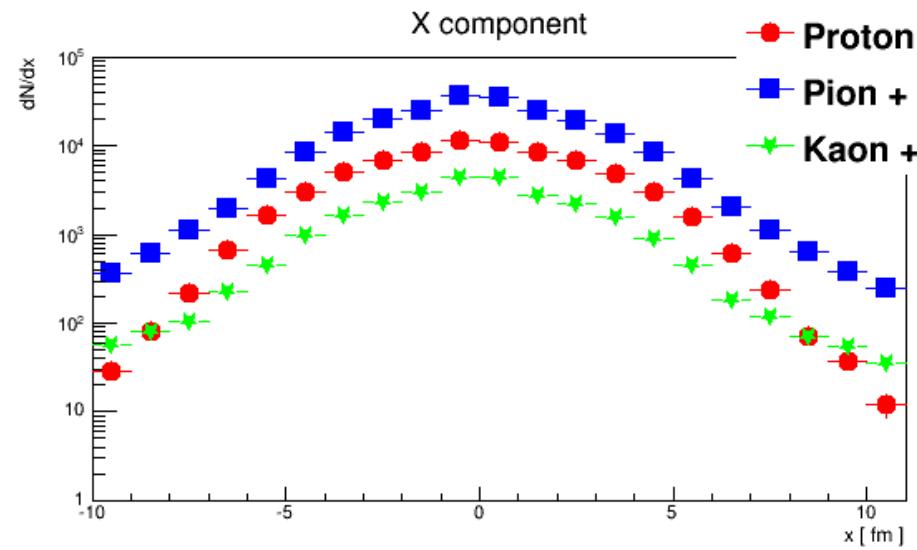
Space four-vector

Particles with negative charge @ 11.5GeV



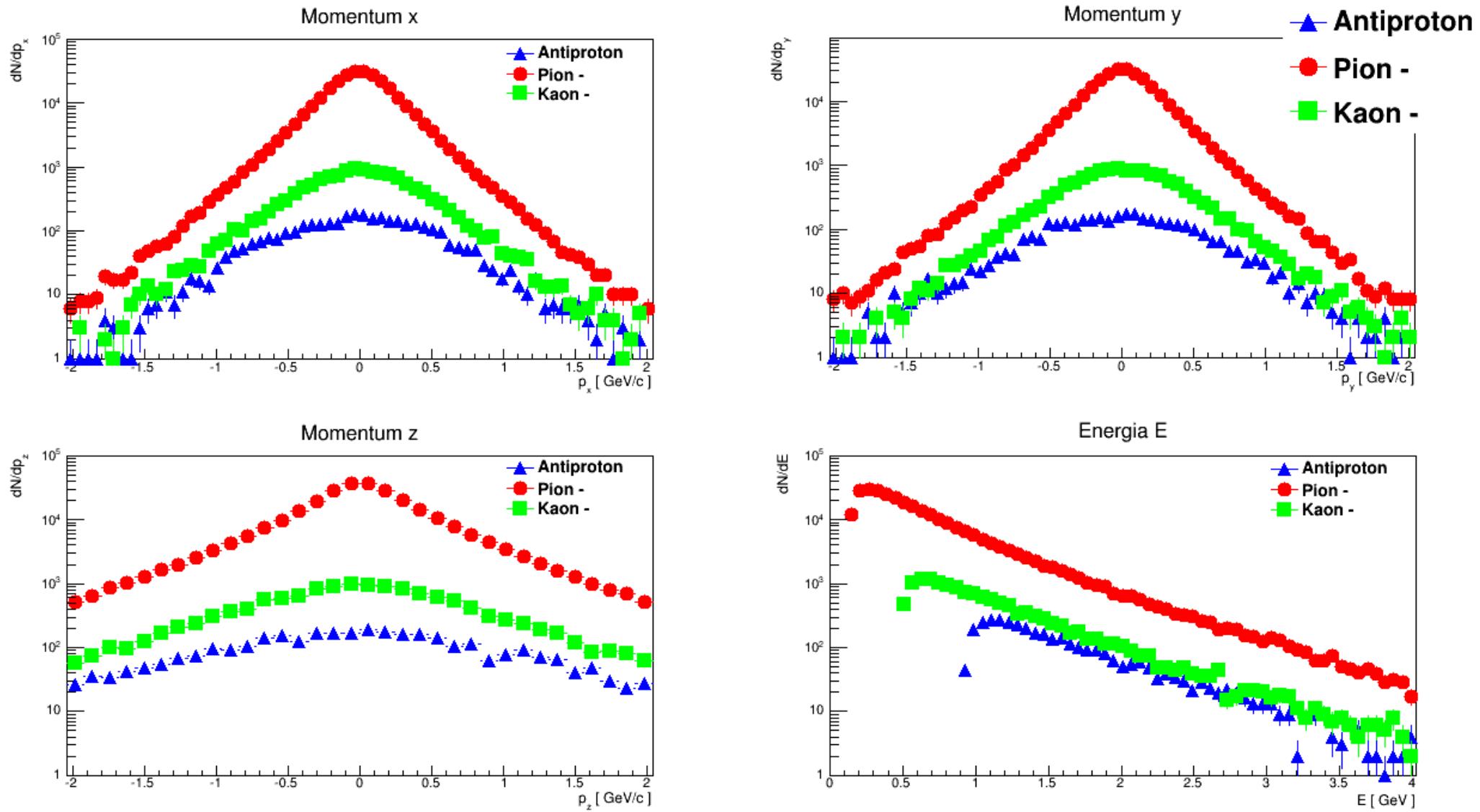
Space four-vector

Particles with positive charge @ 11.5GeV



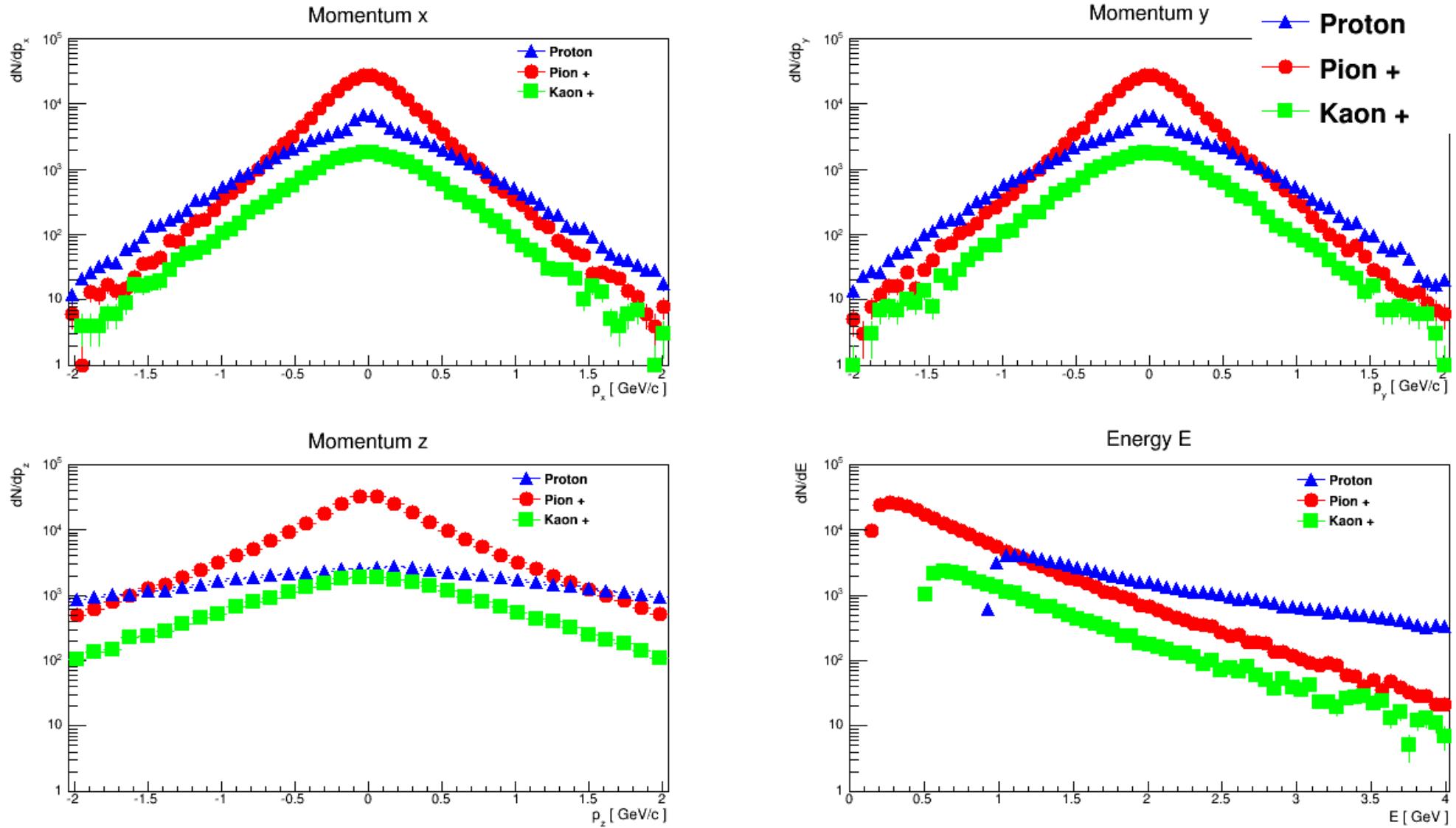
Momentum four-vector

Particles with negative charge @ 11.5GeV

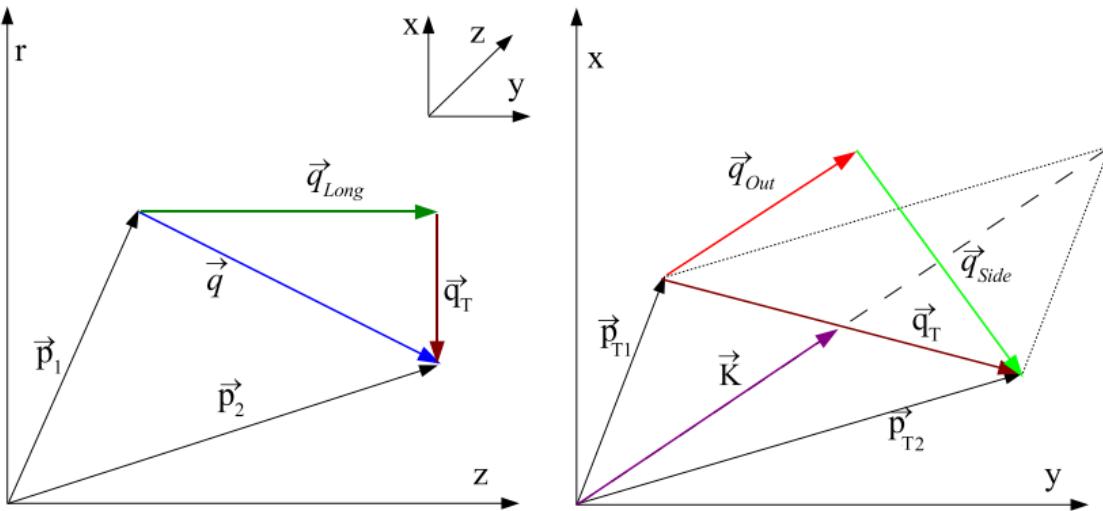


Momentum four-vector

Particles with positive charge @ 11.5GeV



Correlation functions



$$\mathbf{q} = \mathbf{p}_1 - \mathbf{p}_2$$

$\mathbf{k}^* = \mathbf{p}_1 = -\mathbf{p}_2$ (relative pair momentum, calculated in the center of pair mass)

\mathbf{z} – beam direction
 \mathbf{r} – radius

\mathbf{p}_1 and \mathbf{p}_2 – 1 and 2 particle momentum

Correlation function $C(\mathbf{p}_1, \mathbf{p}_2) = \frac{P_2(\mathbf{p}_1, \mathbf{p}_2)}{P_1(\mathbf{p}_1)P_1(\mathbf{p}_2)}$

$P_2(\mathbf{p}_1, \mathbf{p}_2)$ – the probability of the simultaneous two particles emission with momentum \mathbf{p}_1 and \mathbf{p}_2

$P_1(\mathbf{p}_1), P_1(\mathbf{p}_2)$ – the probability of the particle emission with momentum \mathbf{p}_1 or \mathbf{p}_2

Space-time asymmetry

$\cos(\Psi) > 0$

Catching up

Long time of
effective
interaction.

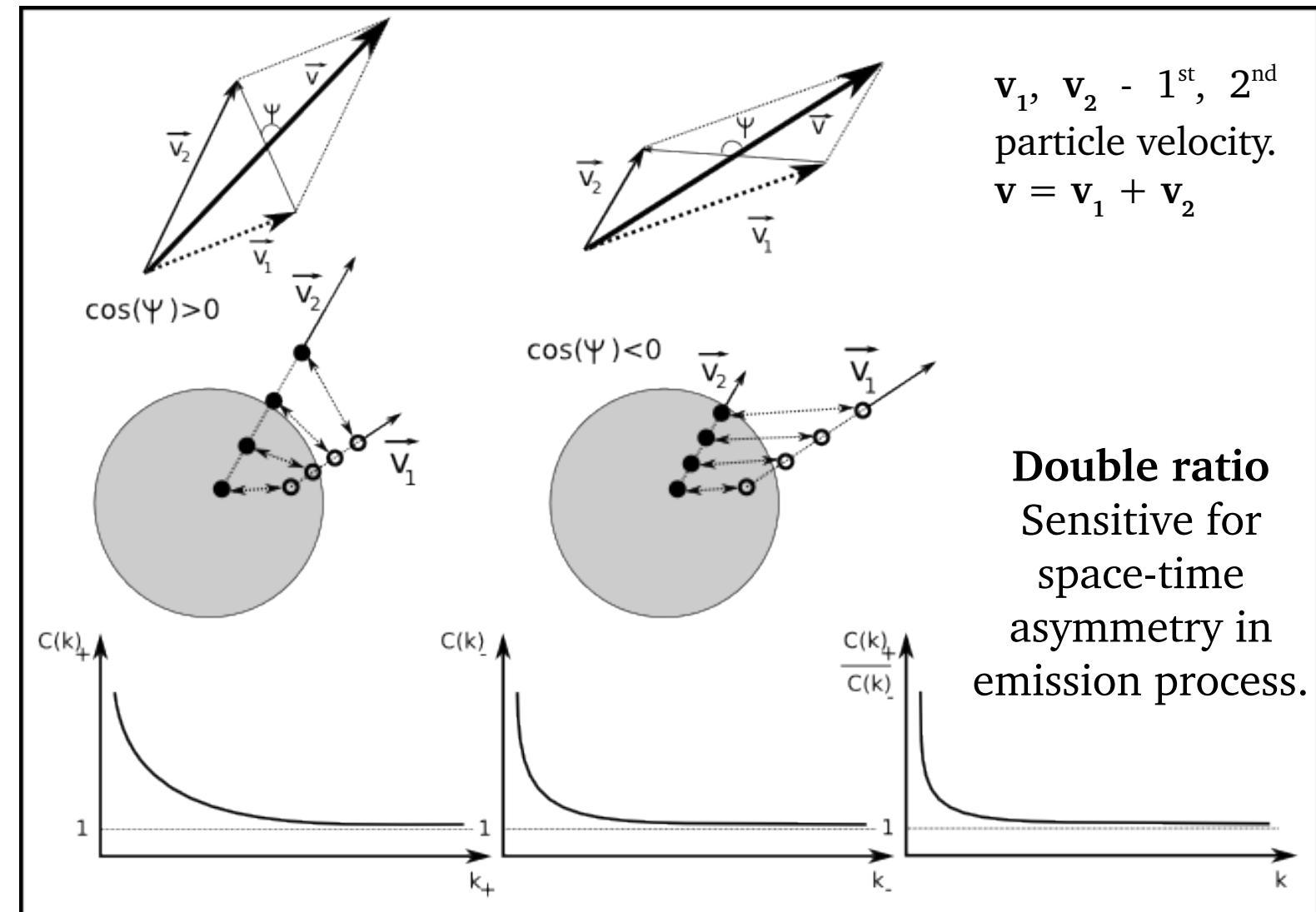
Strong
correlation.

$\cos(\Psi) < 0$

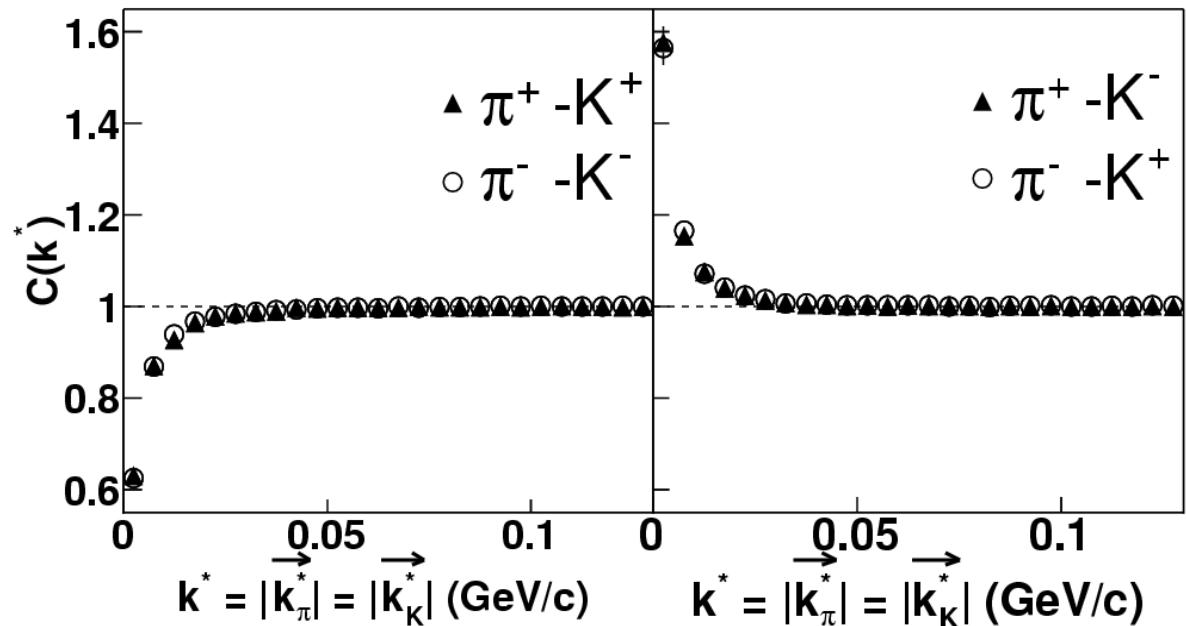
Run away

Short time of
effective
interaction.

Weak
correlation.

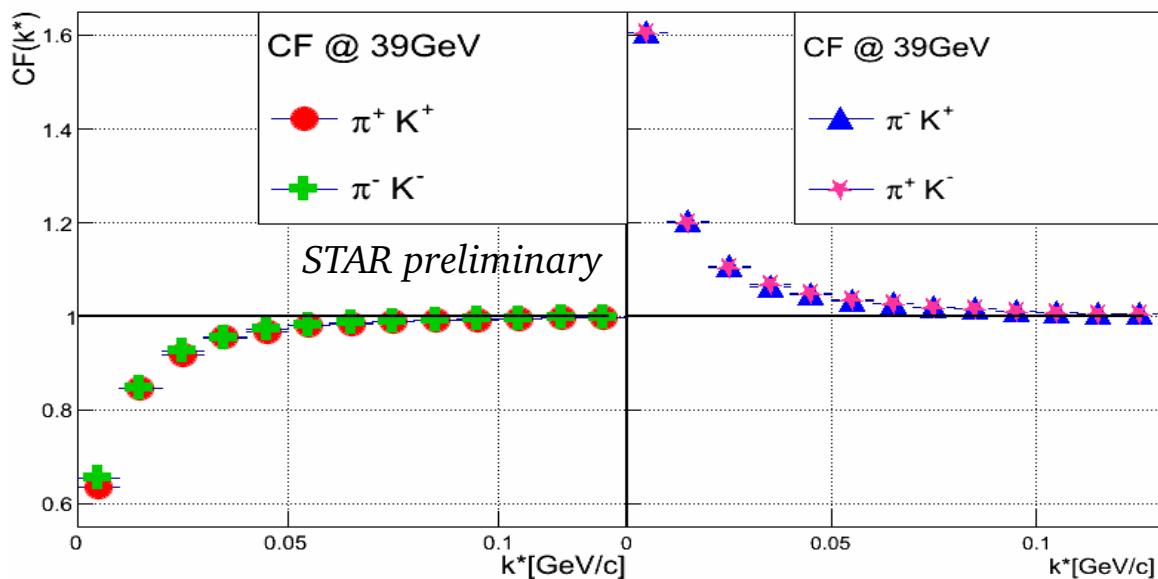


STAR data



AuAu collision at $\sqrt{s_{NN}} = 130\text{GeV}$

Phys. Rev. Lett. 91 (2003) 262302



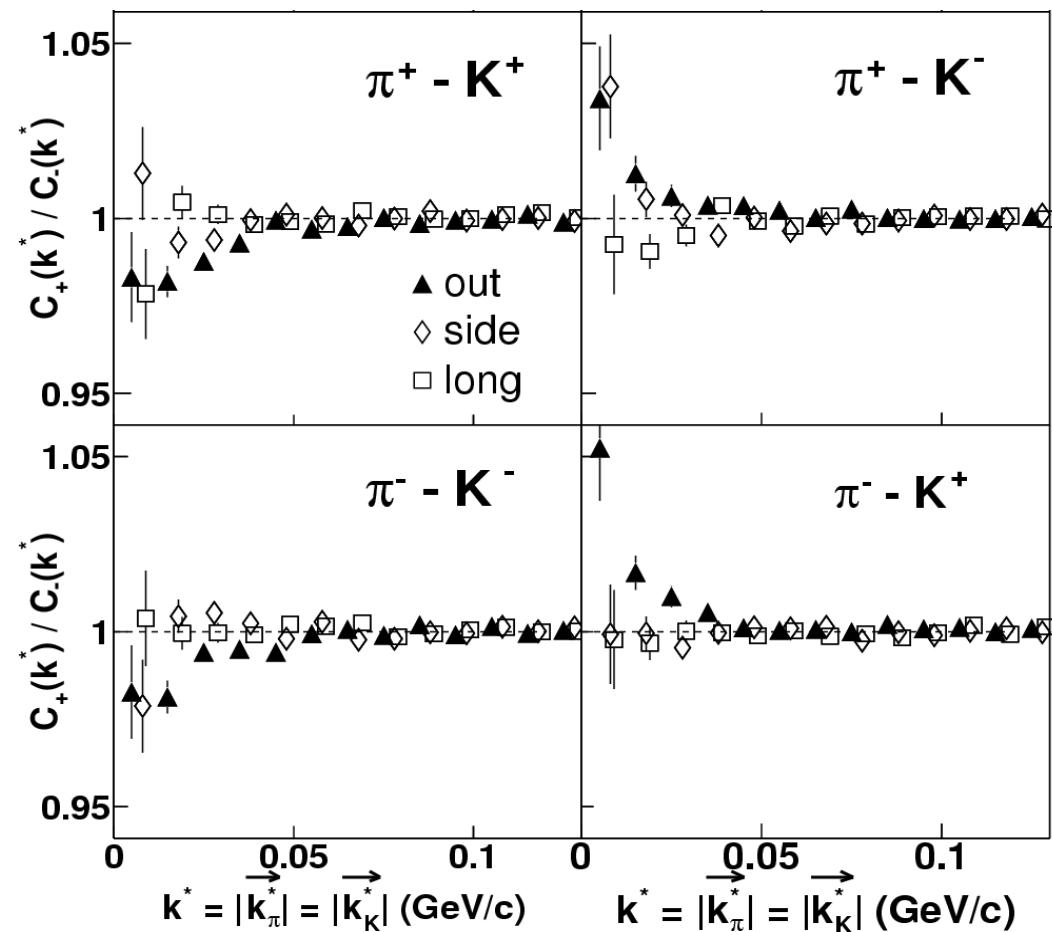
AuAu collision at $\sqrt{s_{NN}} = 39\text{GeV}$

IX WPCF 2013, Acireale (Catania), Italy

STAR data

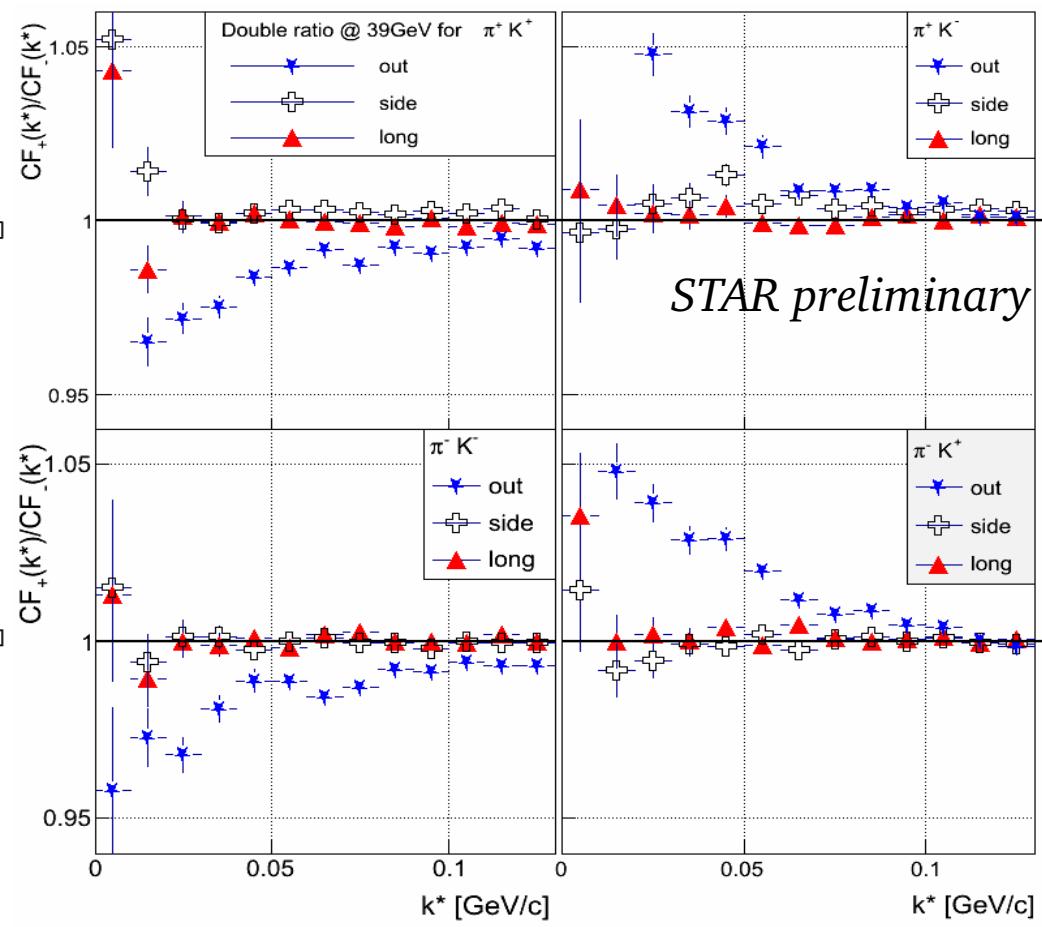
AuAu collision at $\sqrt{s_{NN}} = 130\text{GeV}$

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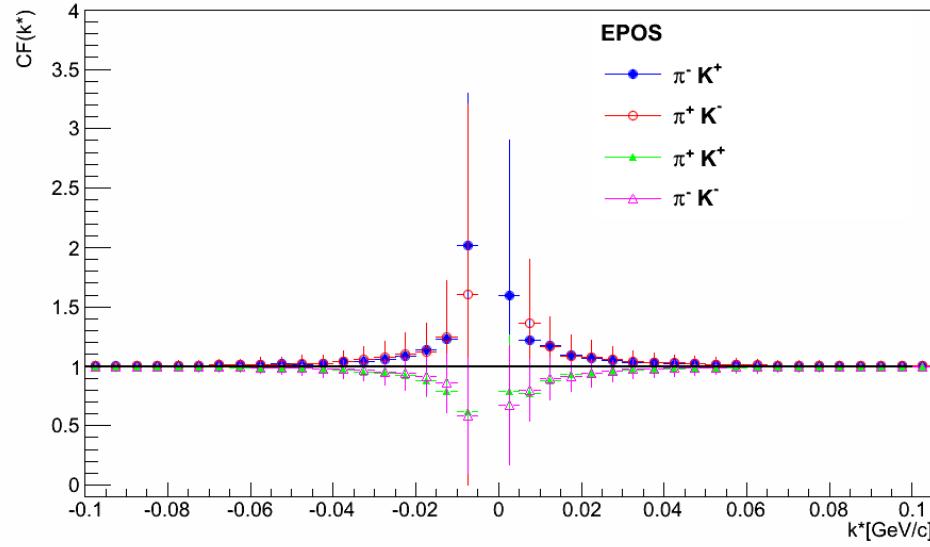
AuAu collision at $\sqrt{s_{NN}} = 39\text{GeV}$

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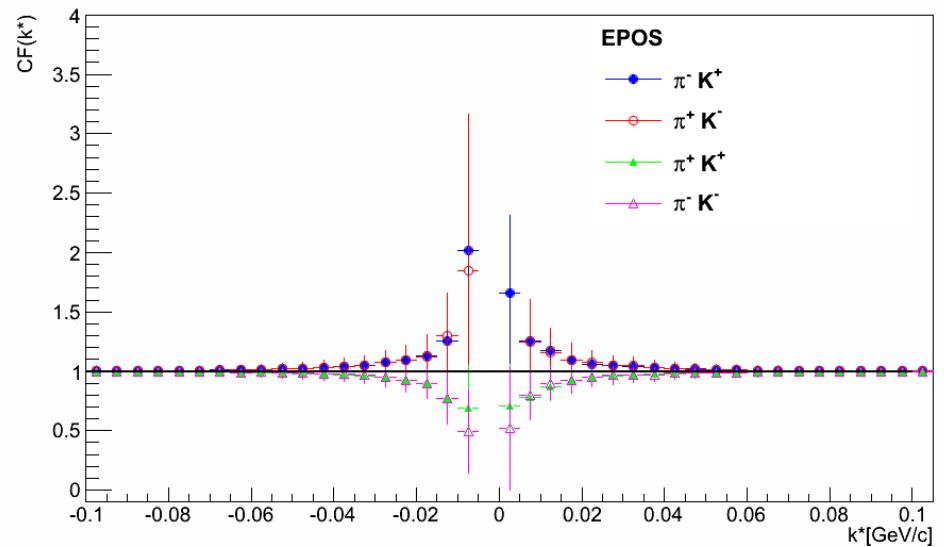


Results from EPOS

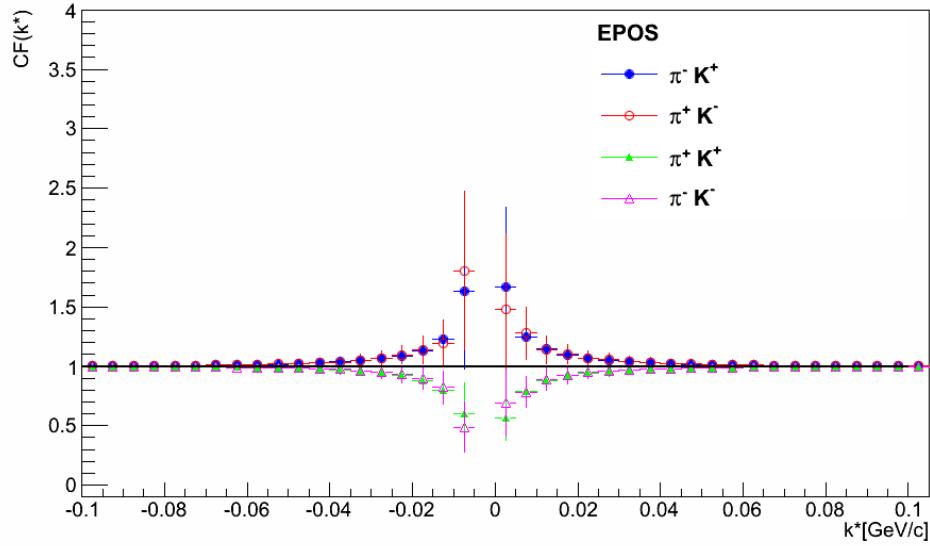
Correlation function @ 7.7GeV



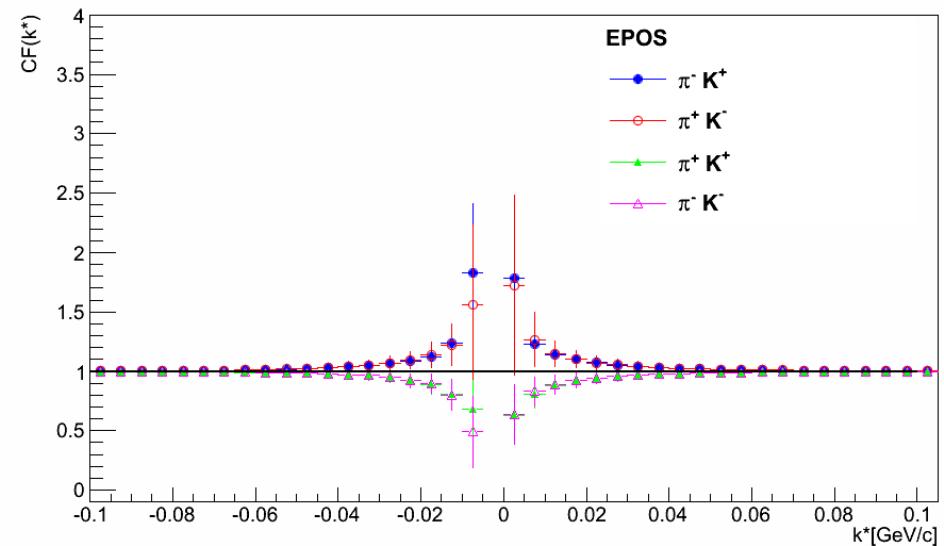
Correlation function @ 11.5GeV



Correlation function @ 19.6GeV

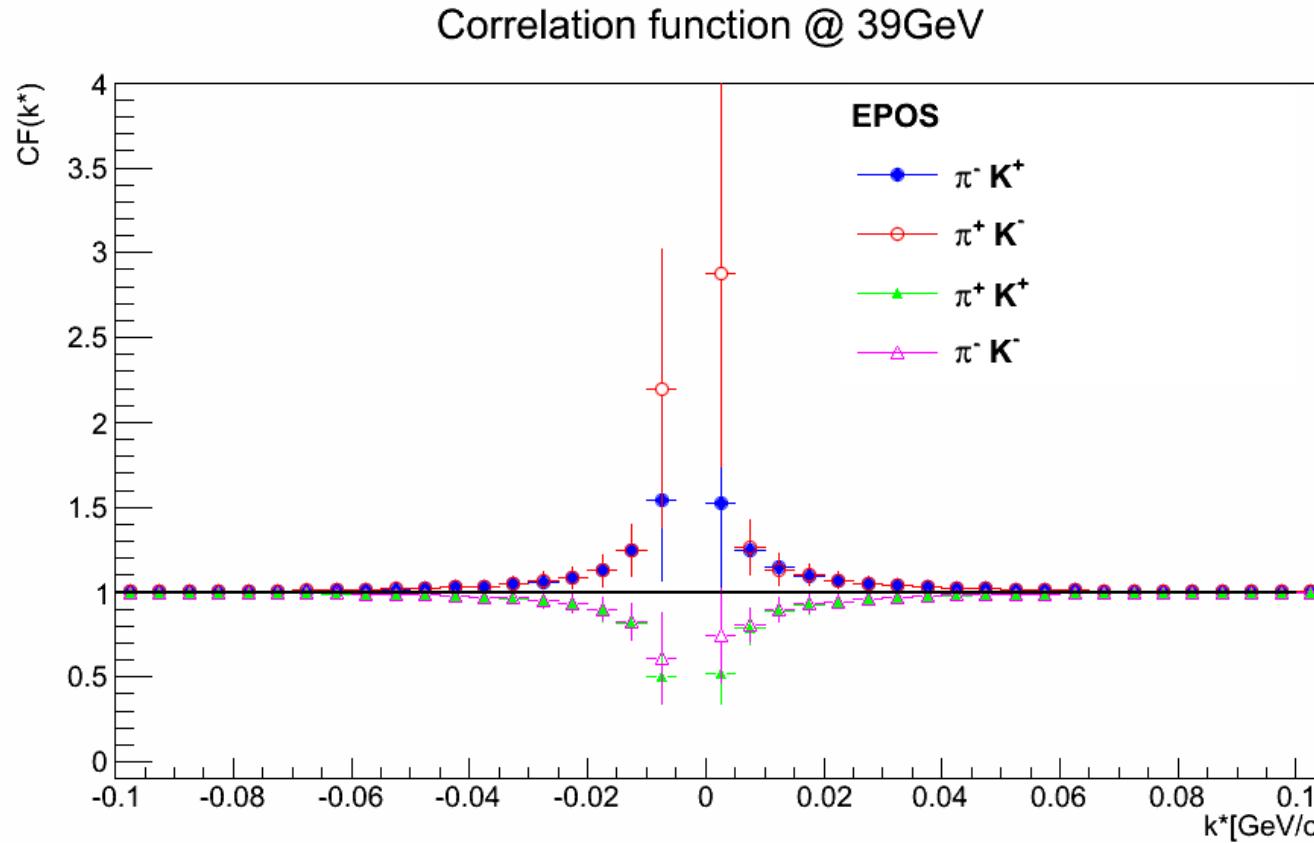


Correlation function @ 27GeV

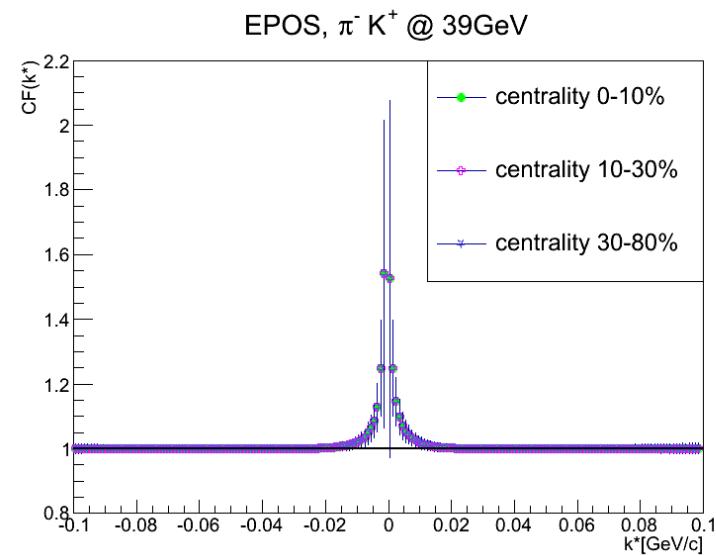
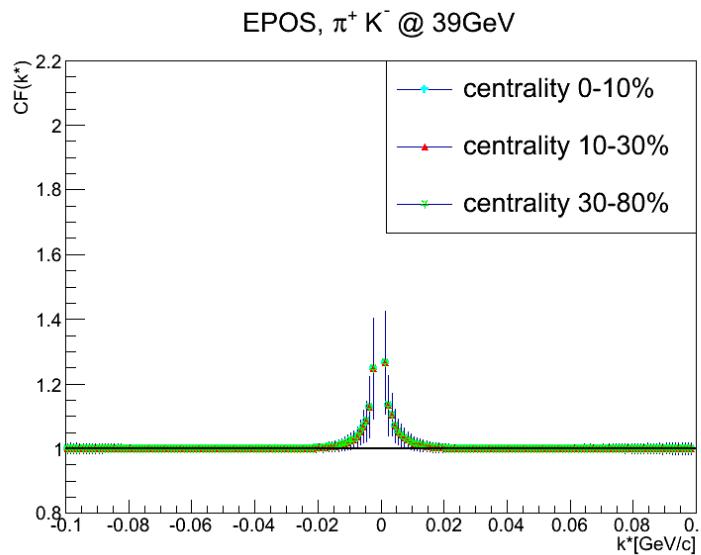
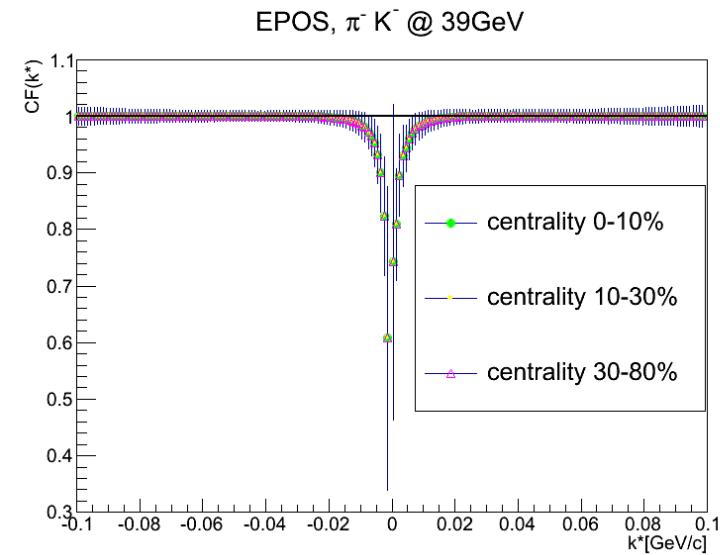
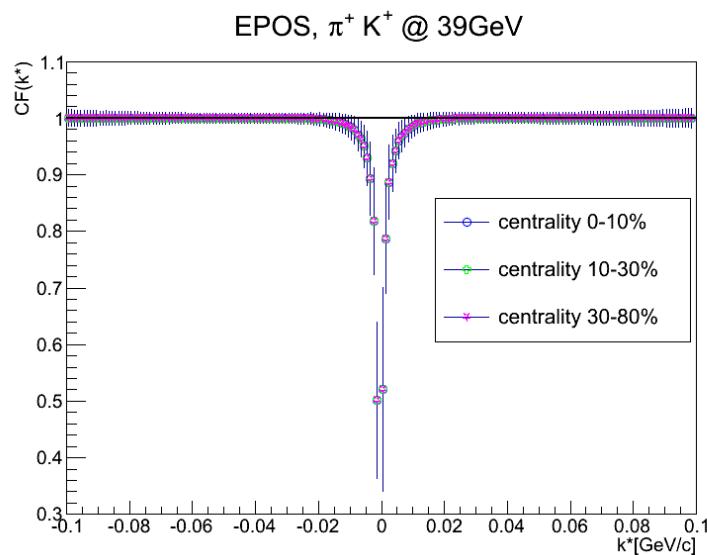


Results from EPOS

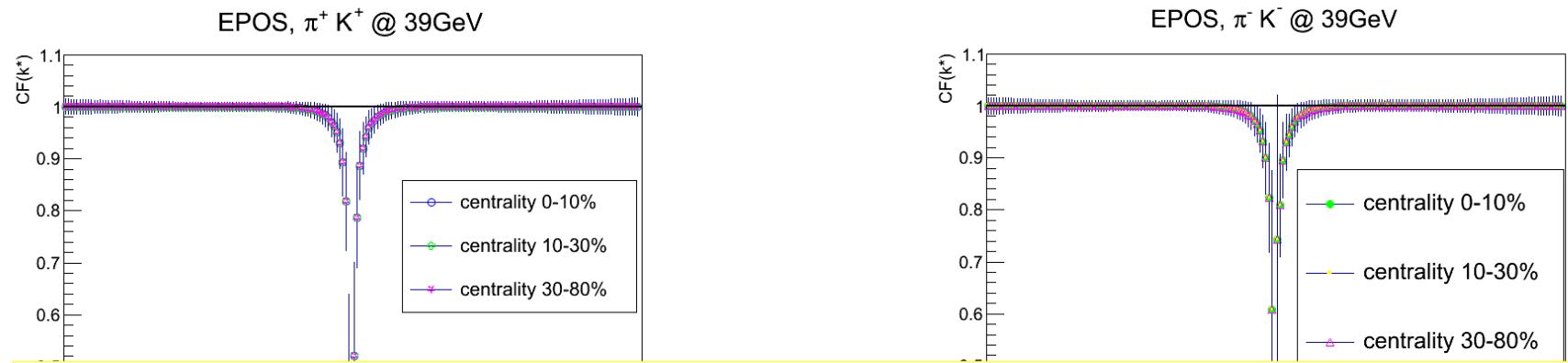
- › The shape of the functions is determined by Coulomb force.
- › In this model there is no big difference between correlation functions of the same sign systems (Pion - Kaon - and Pion + Kaon +) and for correlation functions for different sign systems (Pion - Kaon + and Pion + Kaon -).



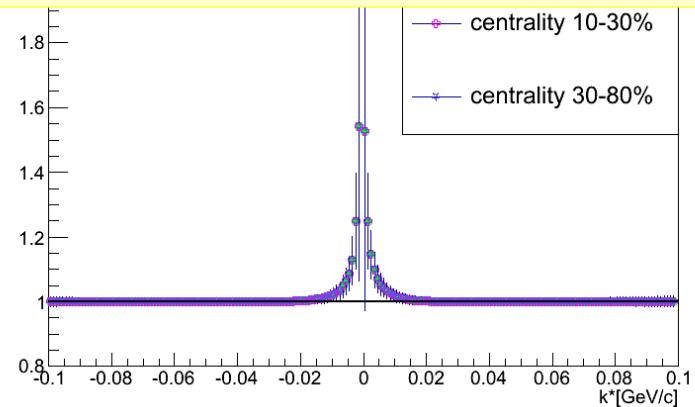
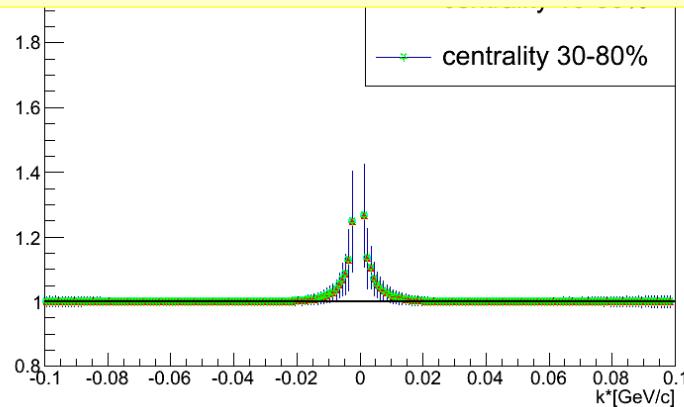
Centrality dependence



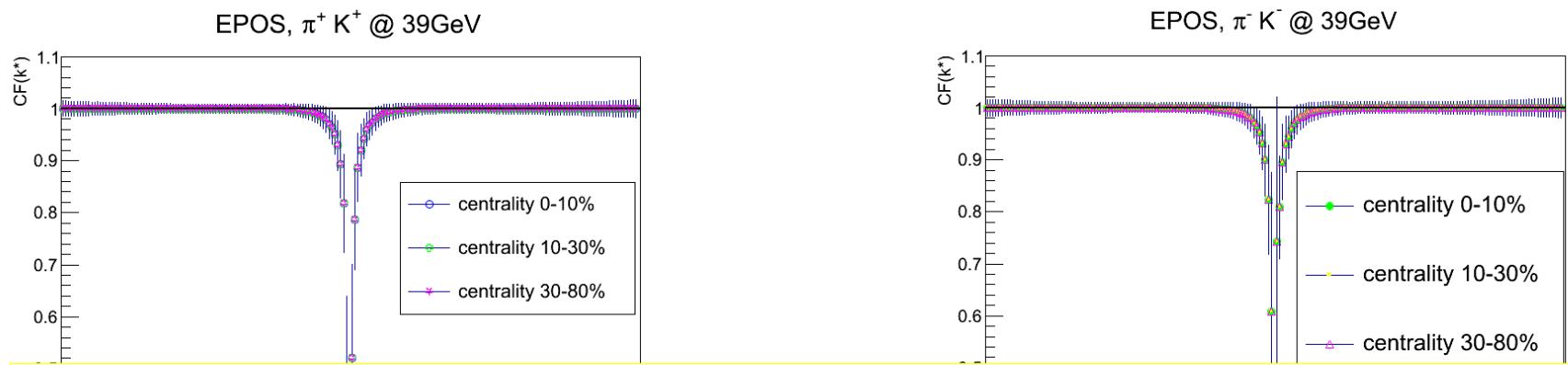
Centrality dependence



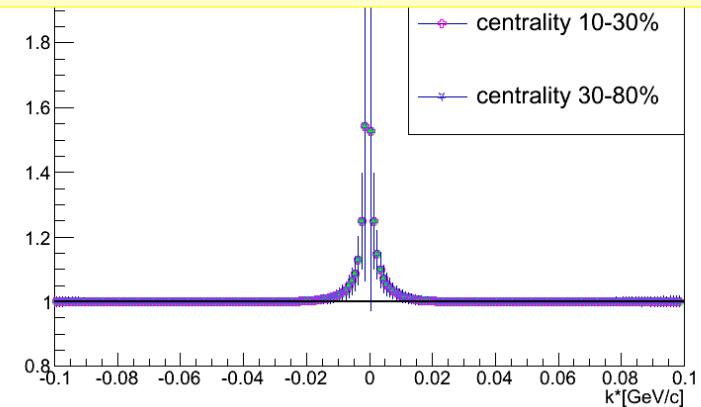
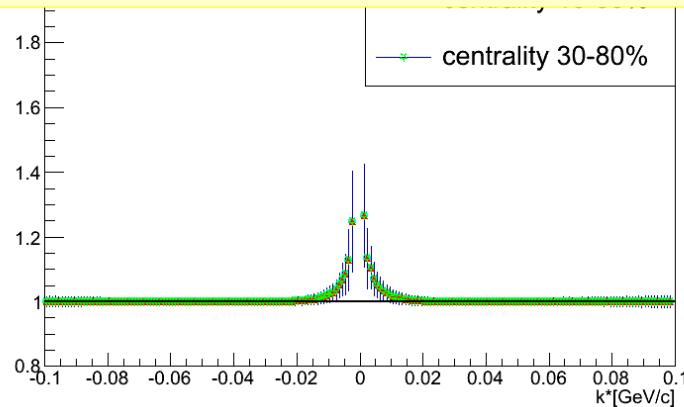
- › No association between the strength of interaction and collision centrality for energy 7.7GeV, 11.5GeV, 19.6GeV, 27GeV and 39GeV.



Centrality dependence

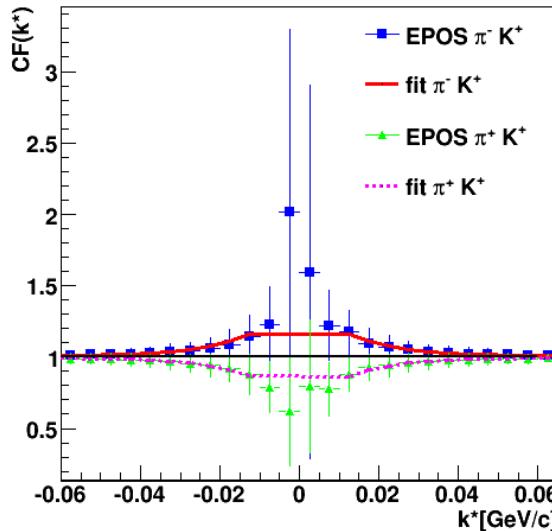


- › No association between the strength of interaction and collision centrality for energy 7.7GeV, 11.5GeV, 19.6GeV, 27GeV and 39GeV.

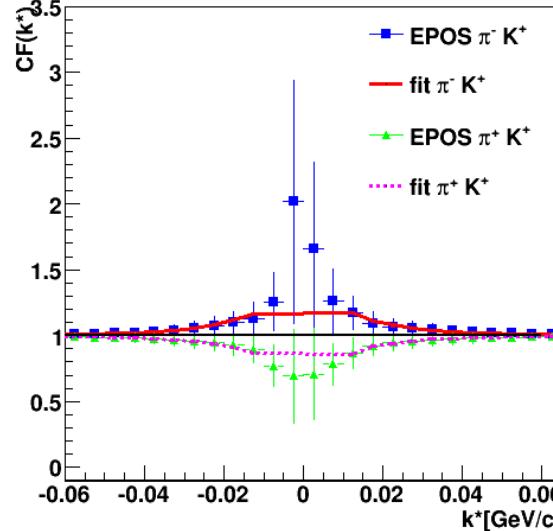


Fits

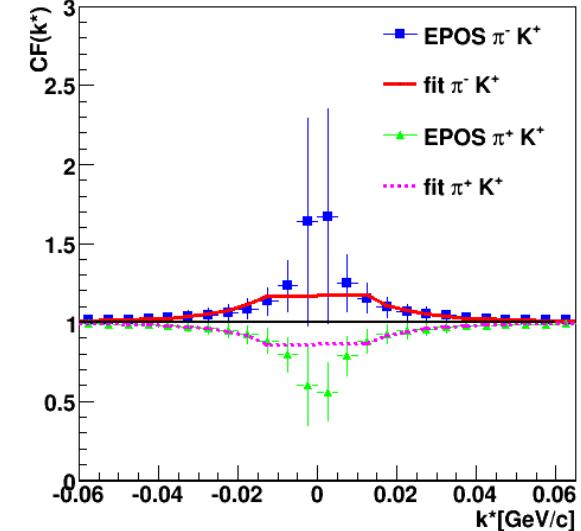
Correlation function with fit @ 7.7GeV



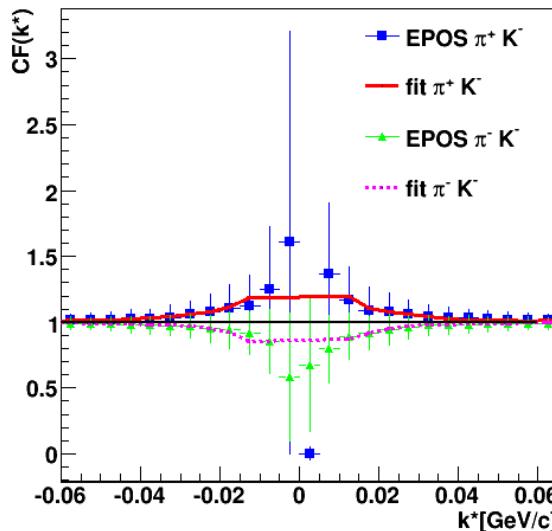
Correlation function with fit @ 11.5GeV



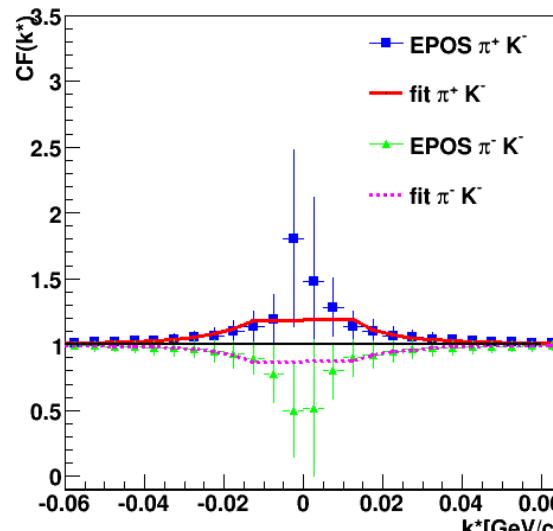
Correlation function with fit @ 19.6GeV



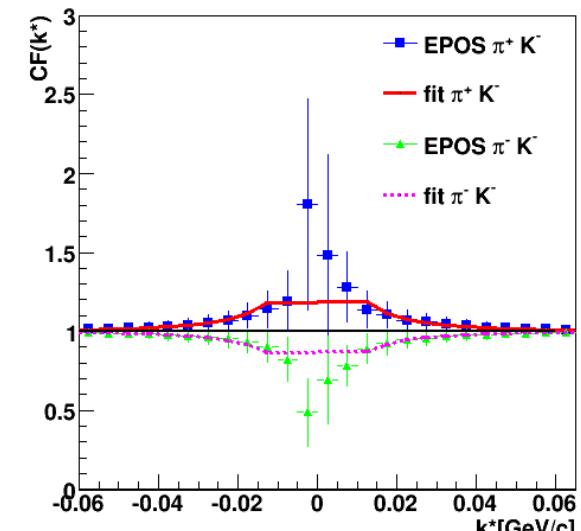
Correlation function with fit @ 7.7GeV



Correlation function with fit @ 11.5GeV

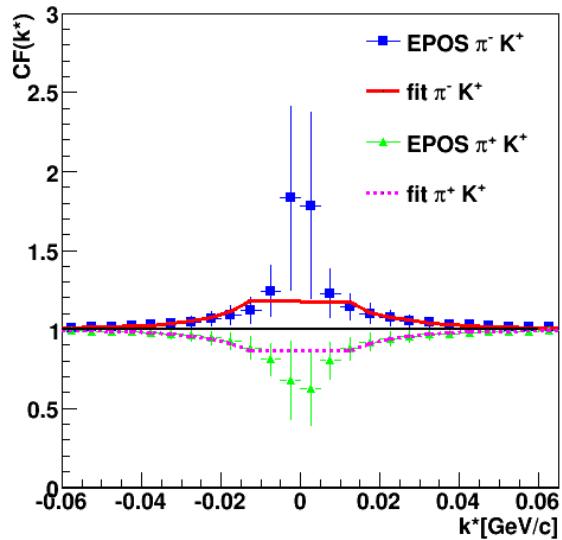


Correlation function with fit @ 19.6GeV

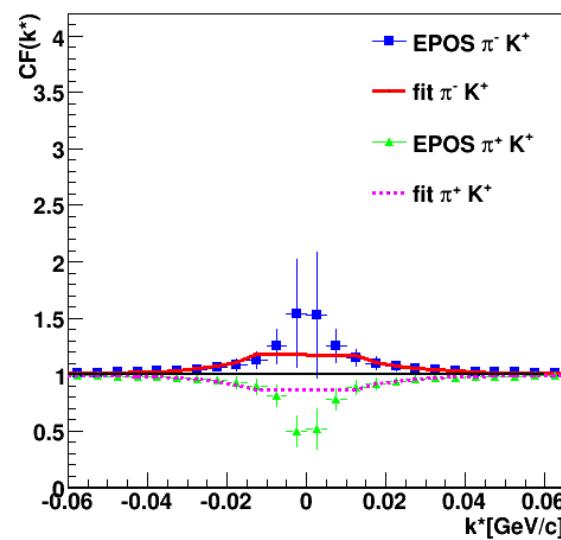


Fits

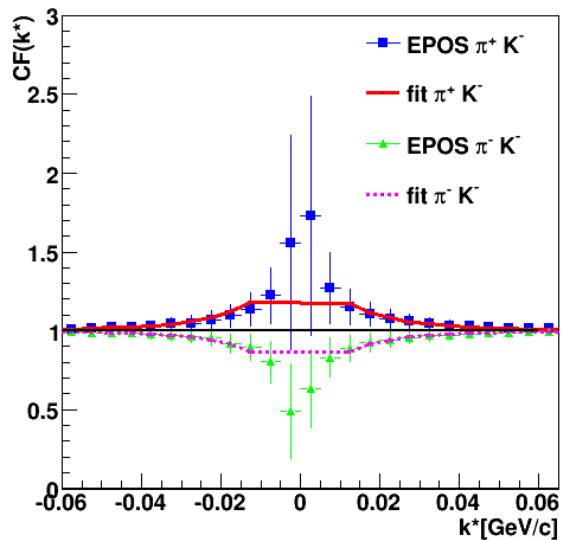
Correlation function with fit @ 27GeV



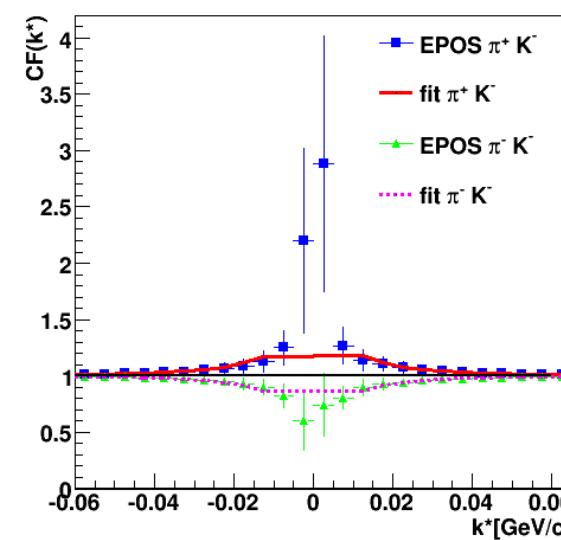
Correlation function with fit @ 39GeV



Correlation function with fit @ 27GeV

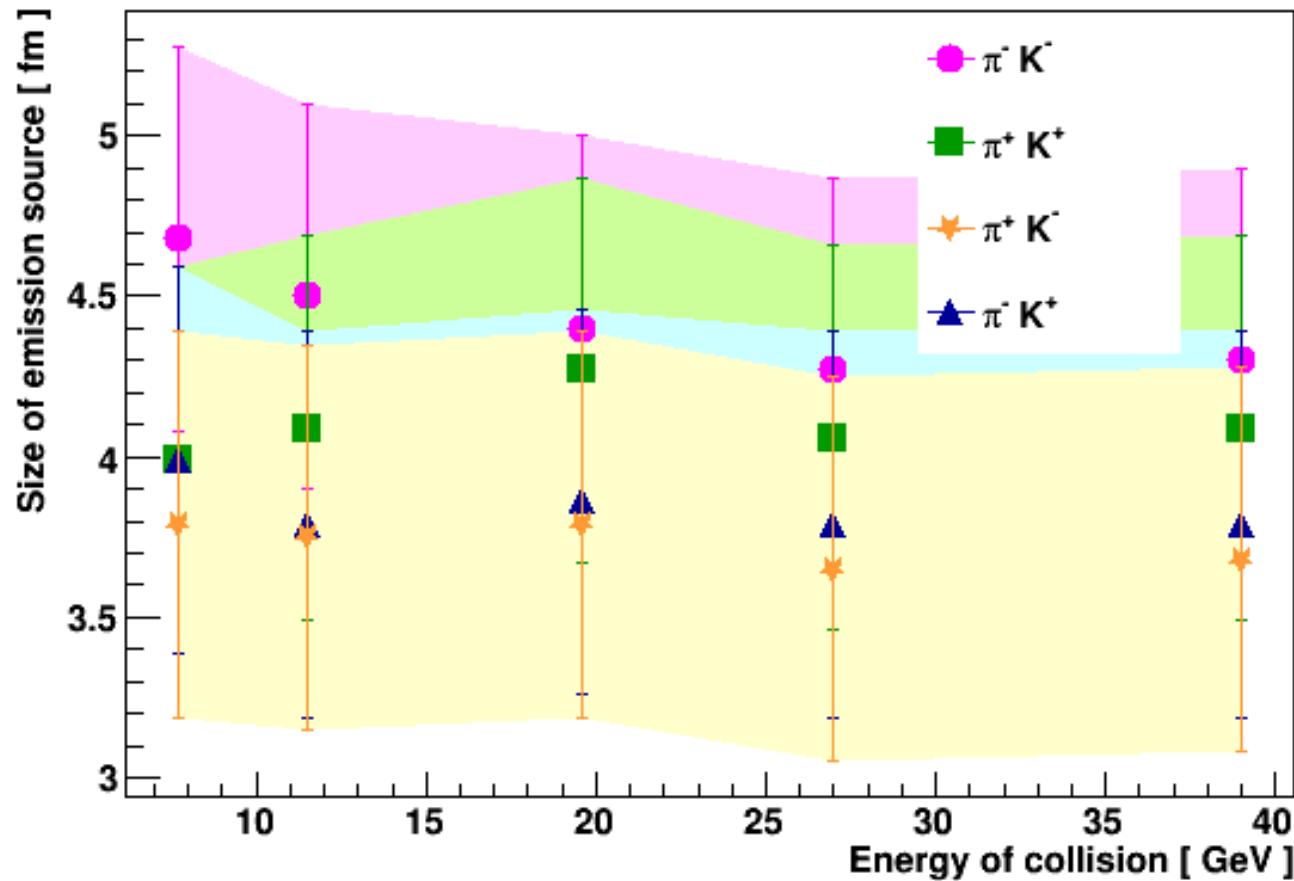


Correlation function with fit @ 39GeV

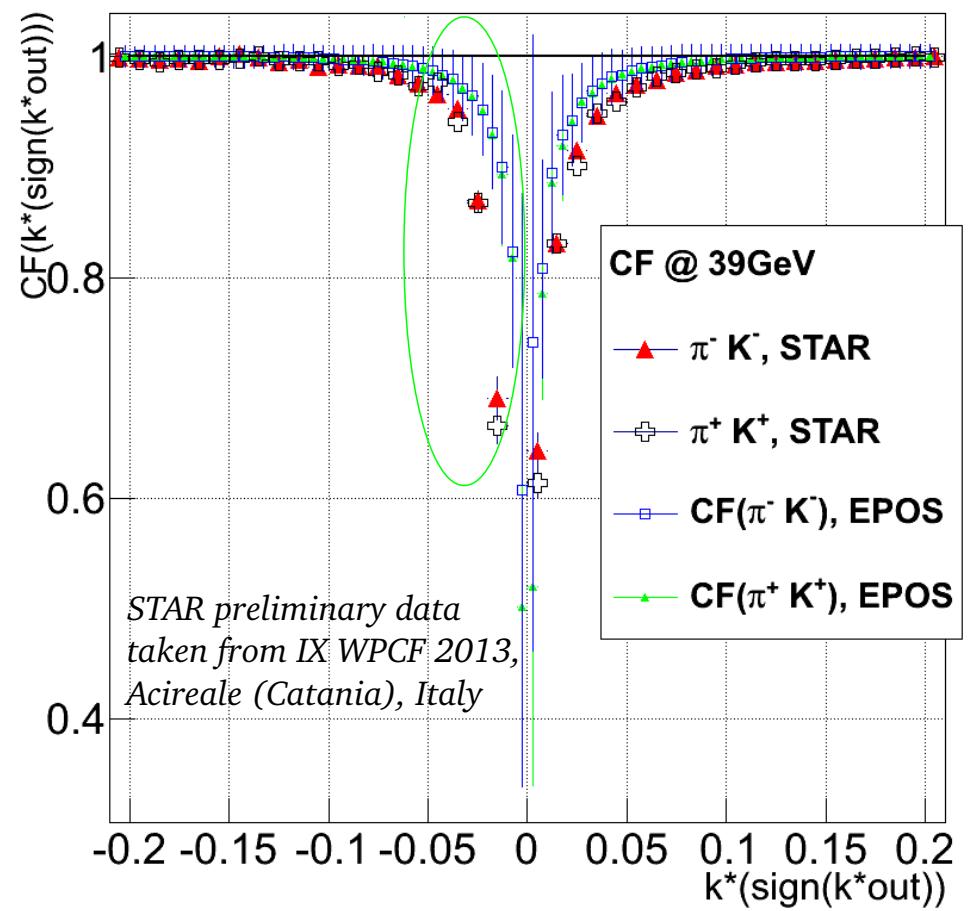
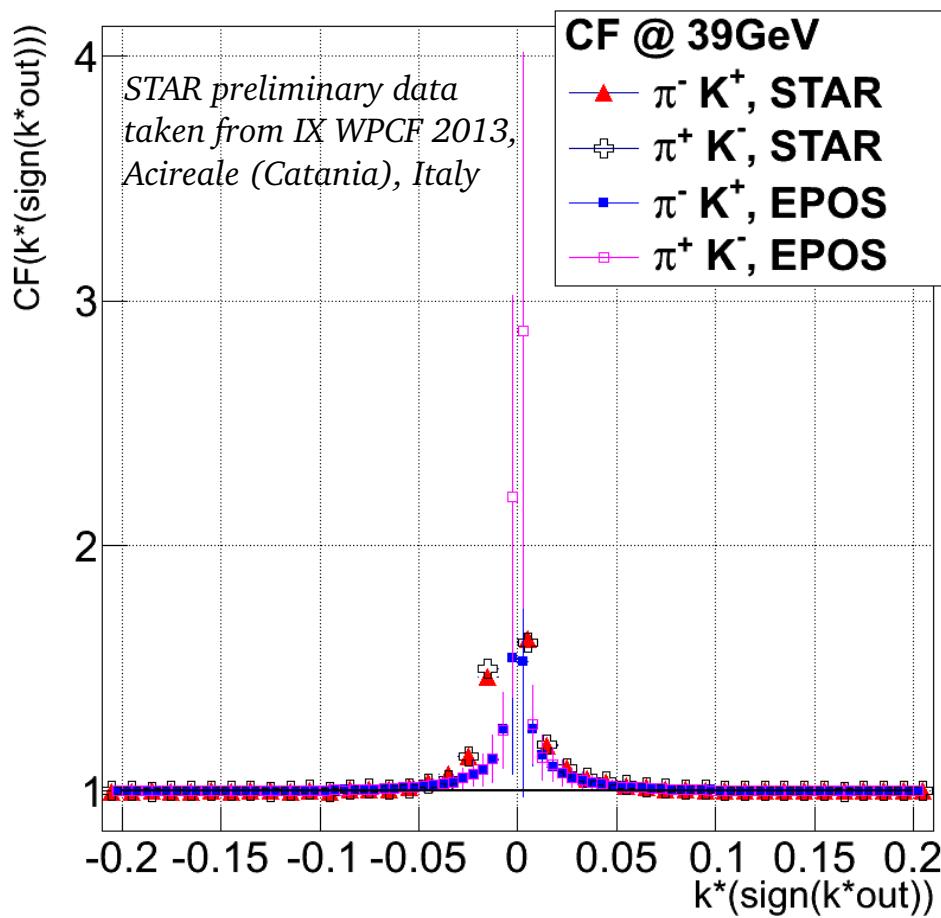


Sizes calculated from correlation functions

- › We do not observe energy dependence in source sizes in EPOS model.

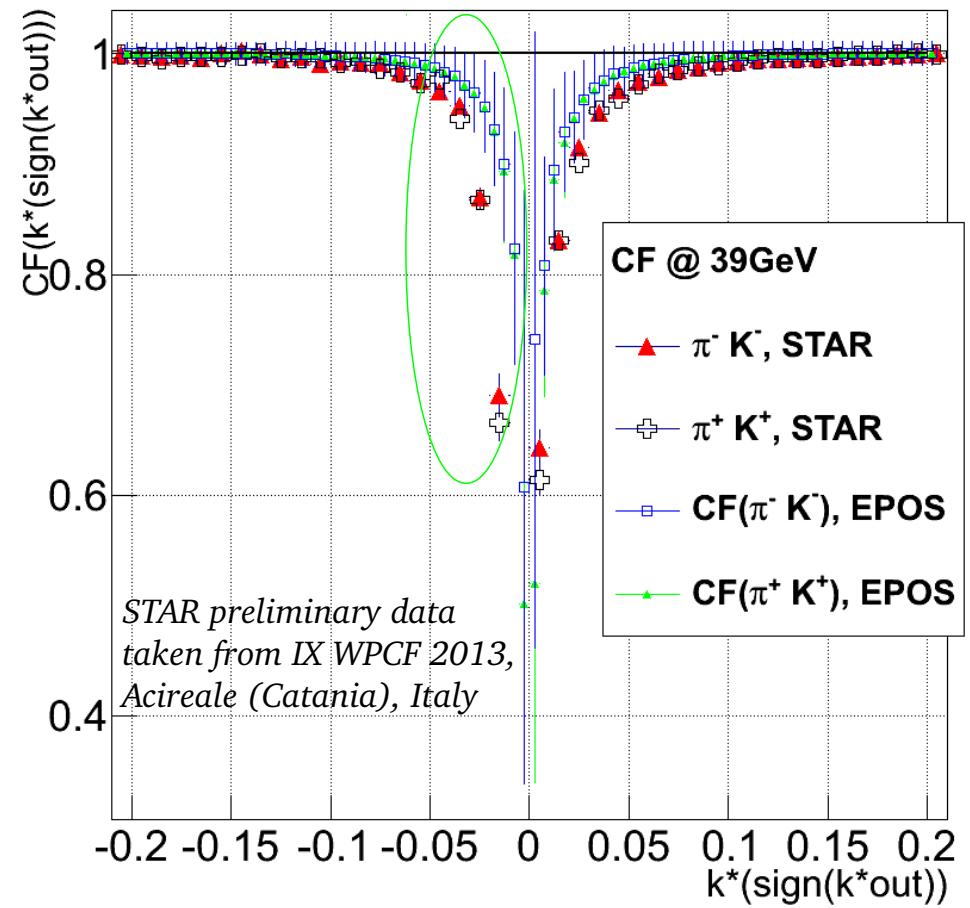
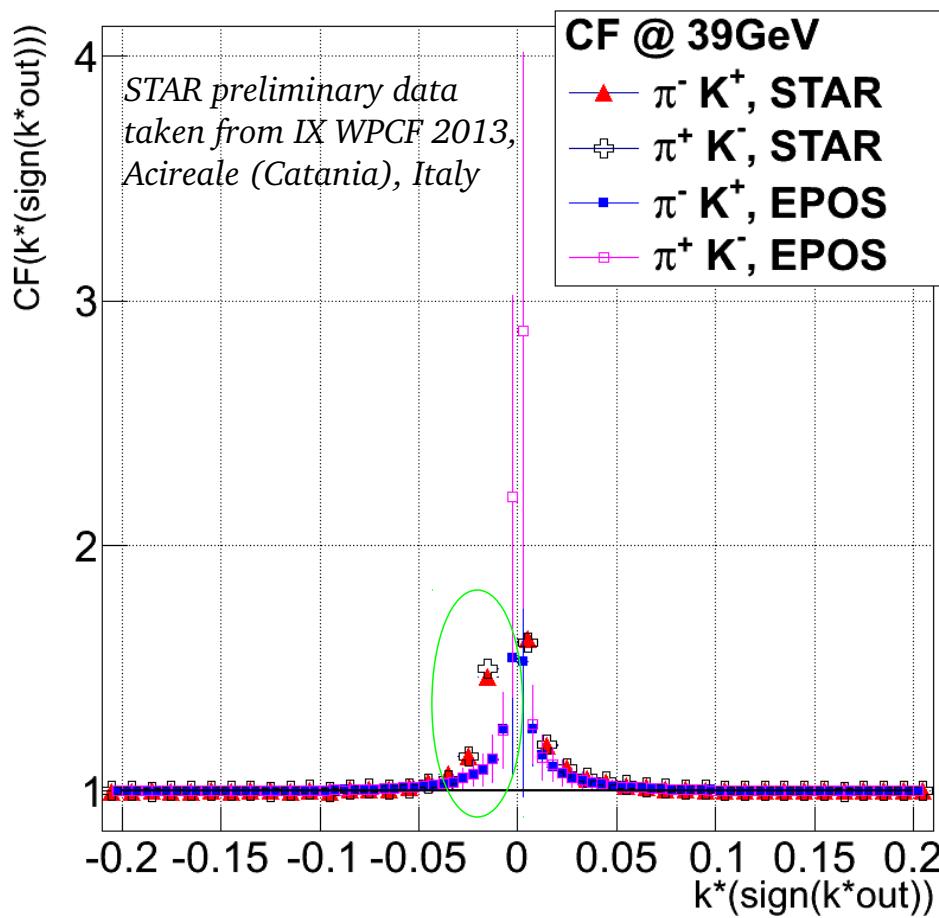


STAR and EPOS data for 39 GeV



STAR and EPOS data for 39 GeV

- › We observe difference between STAR and EPOS data



Summary

- › Transverse momentum spectra results for charged particles from EPOS model are between STAR p_T spectra for the most central collisions (0-5%) and peripheral (60-80%).
- › There is no big difference in EPOS model between correlation functions of the same sign particles systems (Pion - Kaon - and Pion + Kaon +) and for correlation functions for different sign particles systems (Pion - Kaon + and Pion + Kaon -).
- › No association between the strength of interaction and collision centrality for energy 7.7GeV, 11.5GeV, 19.6GeV, 27GeV and 39GeV.
- › We do not observe energy dependence in source sizes in EPOS model.
- › To compare the STAR results from BES with the model, we need the model with hydrodynamics.

Thank you!

Annotation

1. *New Developments of EPOS 2*; T. Pierog from KIT, Institut fur Kernphysik, Karlsruhe, Germany; Iu. Karpenko from Bogolyubov Institute for Theoretical Physics, Kiev, Ukraine; S. Porteboeuf from University of Clermont-Ferrand, Clermont-Ferrand, France; K. Werner from SUBATECH, University of Nantes – IN2P3/CNRS– EMN, Nantes, France
2. Quark Matter 12-18 August 2012, Proceedings; *R_CP and R_AA Measurements of Identified and Unidentified Charged Particles at High p_T in Au+Au Collisions at 7.7, 11.5, 19.6, 27, 39, and 62.4 GeV in STAR*, Evan Sangaline for the STAR Collaboration